

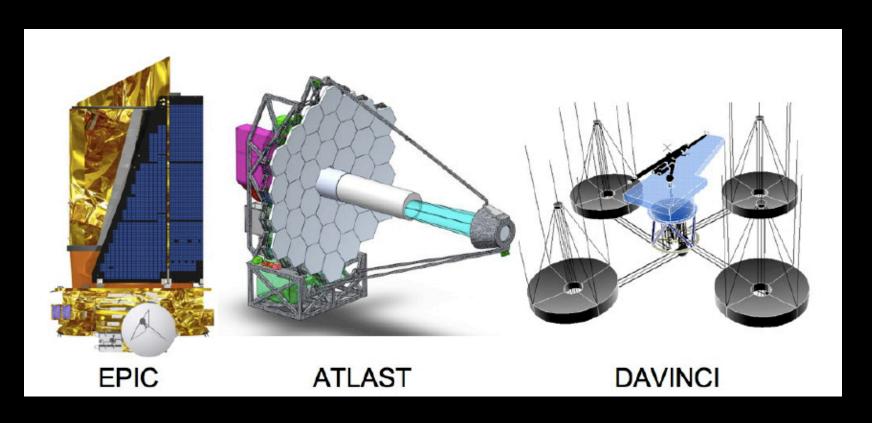
# Visible Nuller Coronagraph Simulations

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(presented by Jagmit Sandhu)

## Visible Nuller Intrument is common to 3 separate NASA ASMC Coronagraph Studies:



- ✓ EPIC : Extrasolar Planetary Imaging Coronagraph (Lyon, Clampin, Woodruff et al. 2010)
- ✓ ATLAST : Advanced Technology for Large aperture Space Telescope (Postman et al. 2010)
- ✓ DAVINCI : Diluted Aperture Visible Nulling Coronagraph Imager (Woodruff, Shao et al. 2010)

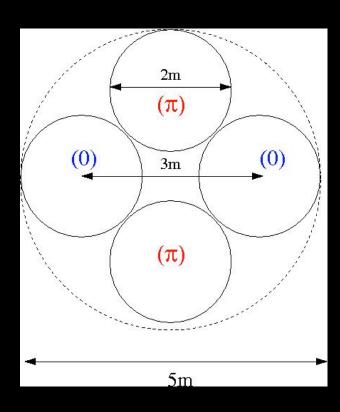


#### General Architecture Studied: 4 Beam Nulling

- ✓ Representative of all 3 ASMC studies (but assumes rigid body pointing)
- ✓ Can simulate arbitrary beam sizes and baseline lengths
- ✓ Simulates speckle fields in 4 x 20% BW filters centered at B,V, R, H
- ✓ Only V band results shown here

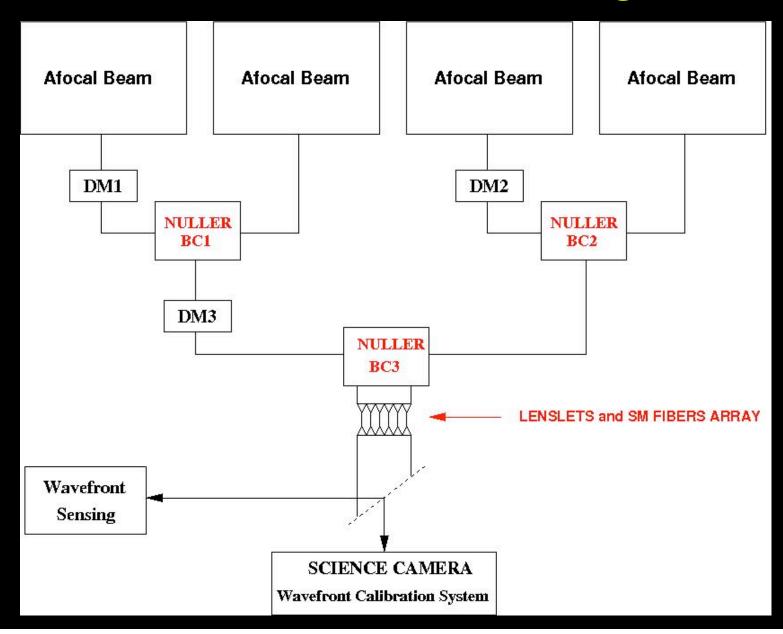
Used individual beam diameter D=2m, baseline B=3m and V band for simulations presented here:

(In the case of a single telescope, the baseline B corresponds to the beam shear, which can be adjusted per star)



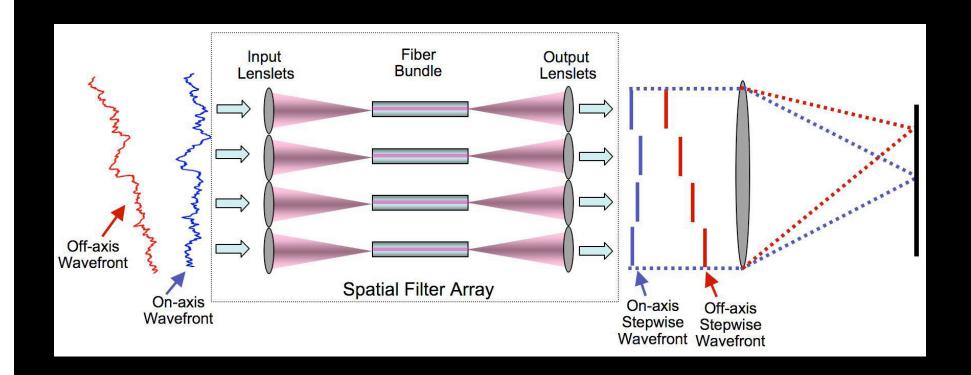


### Instrument Block Diagram





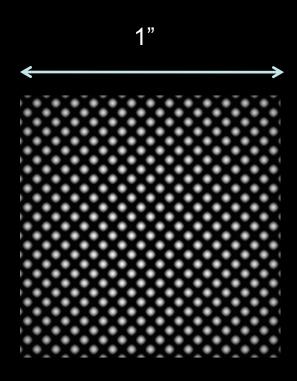
### Zoom on SM Fiber Array



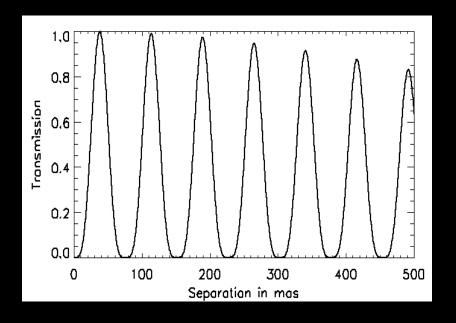
Simulations assume hexagonally shaped Array of 1027 SM Fibers and hexagonally packed lenslets



### **Sky Transmission Pattern**



Sky transmission pattern (tapering due to SM fiber FOV)



Monochromatic transmission at 550nm along X axis Half power point is  $(2/\pi) \lambda/B$  That is 0.85  $\lambda/D$  with D=4m (diameter adopted for Lyot and external occulter designs)



Assumes science integration is made of many DM piston and tip-tilt corrections of finite duration, resulting in:

✓ Finite Wavefront Measurement SNR → Residual rms error on both amplitude and phase for each beam and for each fiber (spatial white noise floor)

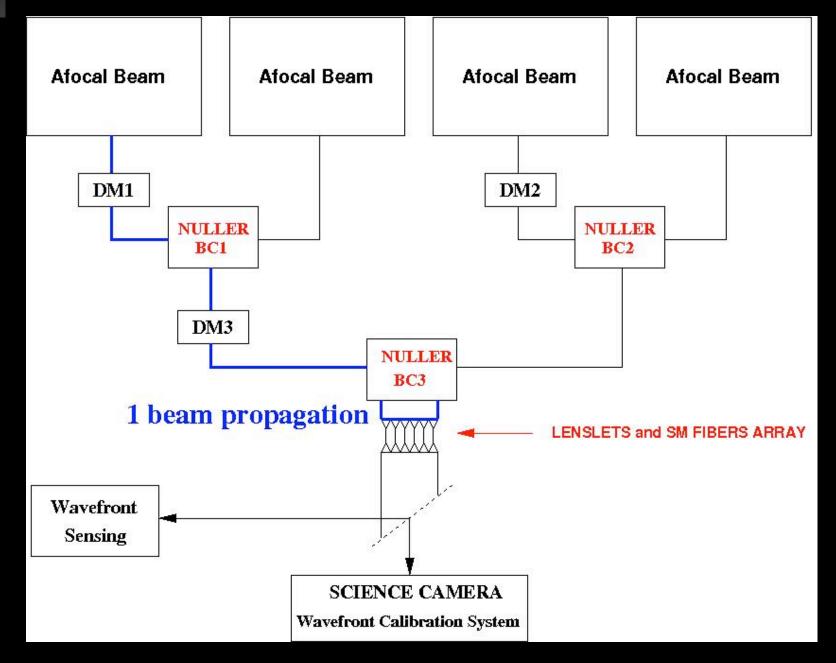


- ✓ Finite Wavefront Measurement SNR → Residual rms error on both amplitude and phase for each beam and for each fiber (spatial white noise floor)
- ✓ Dynamic Wavefront Distortion between Successive Wavefront measurements and Corrections → low order spatial corrugations, primarily at telescope level



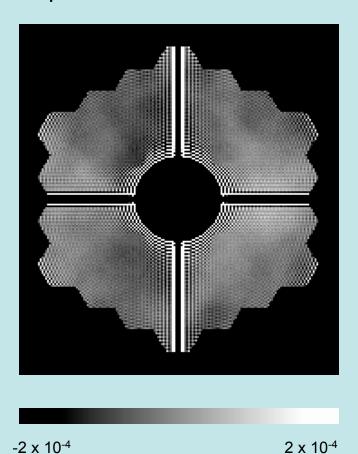
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- ✓ Dynamic Wavefront Distortion between Successive Wavefront measurements and Corrections → low order spatial corrugations, primarily at telescope level
  - ✓ Use "PROPER" Finite Diffraction Code to propagate individual beams (or telescopes) distortions to Fiber Lenslet Array



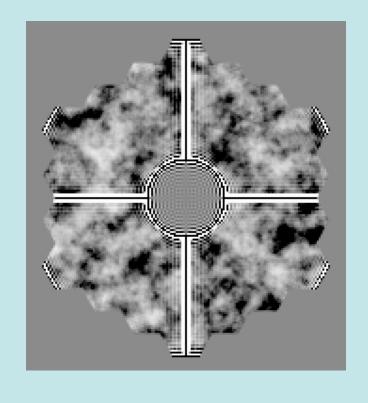


### Individual Telescope (or Beam) Amplitude and Phase maps after propagation to 1st lenslets array

Amplitude Relative Fluctuations



Phase (radians)



2 x 10<sup>-4</sup>

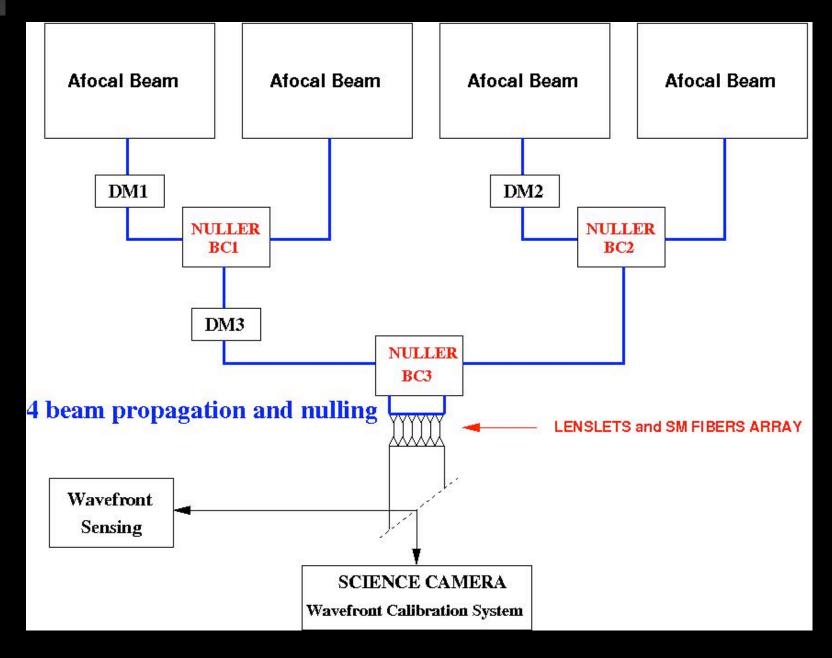
-2 x 10<sup>-4</sup>

Example: 5pm rms uncorrected surface error and 2x10<sup>-5</sup> rms amplitude error (power law PSDs)



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  - ✓ Use "PROPER" Finite Diffraction Code to propagate individual beams
    ✓ (or telescopes) distortions to Fiber Array
- ✓ Compute Nulled Field Amplitude and Phase Distribution before & after injection into each Single-Mode Fiber



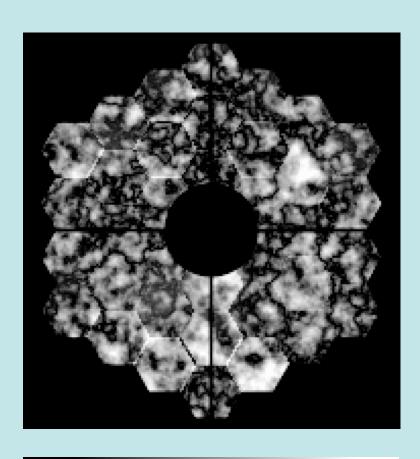


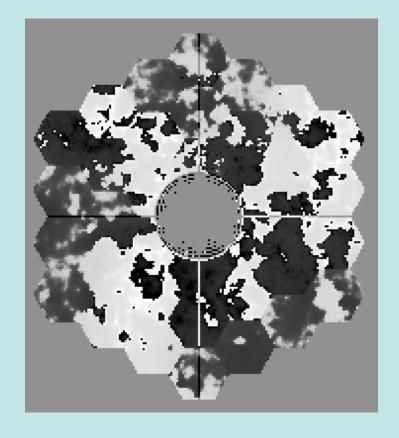


## Nulled Electric Field Amplitude and Phase right after DM correction (before fibers):

Residual Amplitude

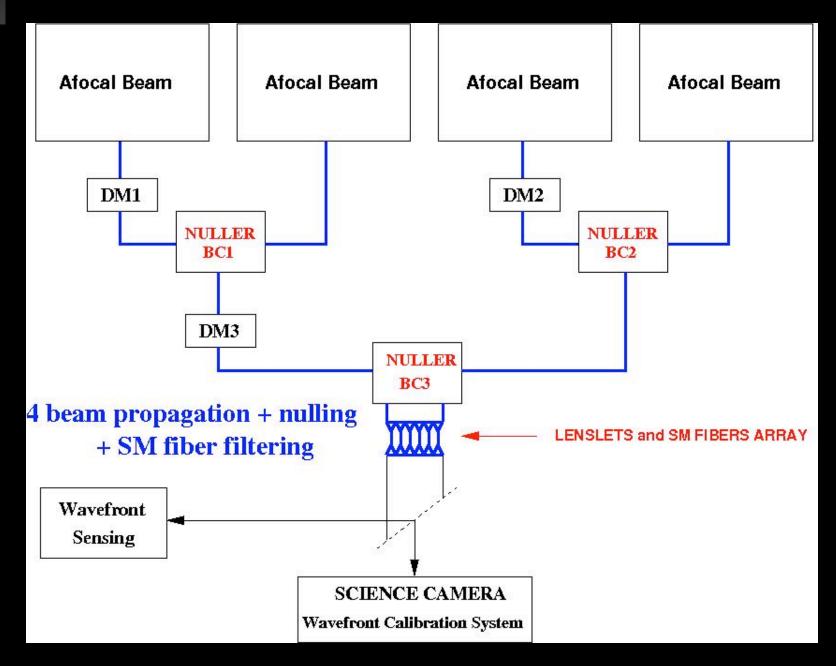
Phase (radians)





0 1.5 x 10<sup>-4</sup> 3 x 10<sup>-4</sup> - $\pi$ 



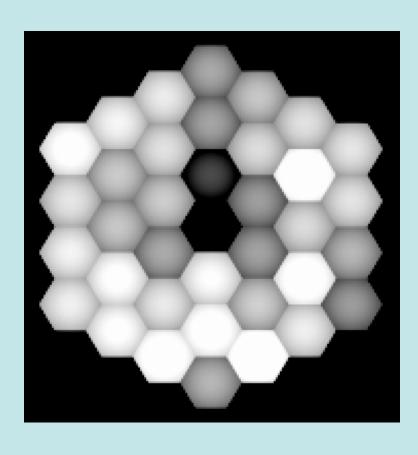




#### Nulled Electric Field amplitude and phase after DM correction and SM fiber filtering

Residual Amplitude

Phase (radians)





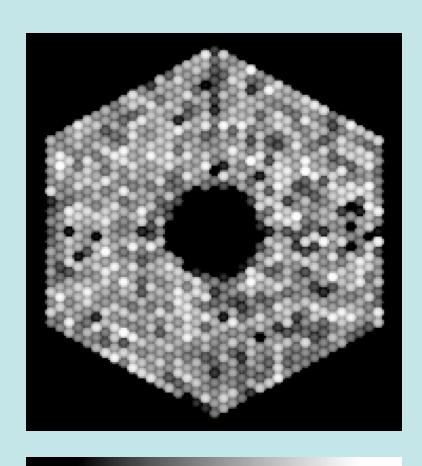
10-6 10-5 10-4

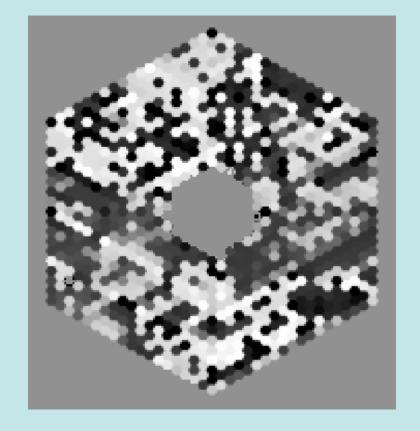


## Nulled Electric Field amplitude and phase after DM correction and SM fiber filtering

Residual Amplitude

Phase (radians)





10<sup>-6</sup> 10<sup>-5</sup> 4 x10<sup>-4</sup>

 $-\pi$ 

 $\pi$ 



- ✓ Finite Wavefront Measurement SNR → Residual rms error on both amplitude and phase for each beam and for each fiber (spatial white noise floor)
- ✓ Dynamic Wavefront Distortion between Successive Wavefront measurements and Corrections → low order spatial corrugations, primarily at telescope level
- ✓ Use "PROPER" Finite Diffraction Code to propagate telescope distortions to Fiber Array
- ✓ Compute Nulled Field Amplitude and Phase Distribution after injection into each single-mode fiber
  - ✓ Repeat for many wavelengths inside a given filter



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  - ✓ Repeat for many wavelengths inside a given filter
    - ✓ Repeat and Average over many instances



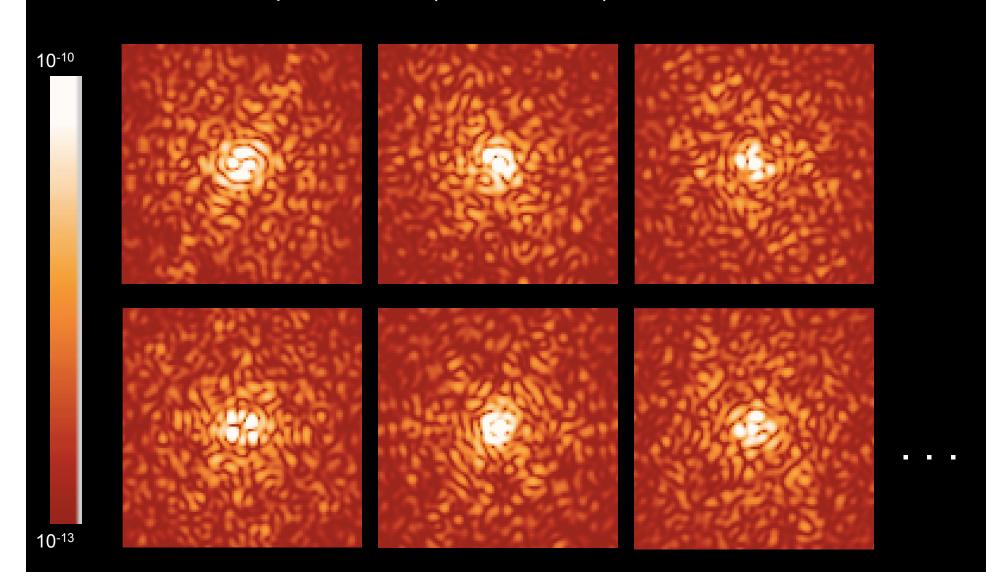
#### Simulations Parameters

- √ 11 wavelengths between 500 nm and 600 nm
- ✓ Segmented DM with 1027 hexagonal actuators (independent piston and tip-tilt controls for each)
- ✓ WFC error (due to Telescope Distortions in between WFC):
  - 0.10 mas rms tip-tilt per axis
  - 10pm total phase rms per beam, with power law PSD
  - 2 10<sup>-5</sup> amplitude rms *per beam*, with power law PSD
- ✓ WF measurement error floor (induced by photon noise during WFS)
  - Residual opd= 10 pm rms per beam & per fiber.
  - Residual amplitude mismatch= 2 10<sup>-5</sup> per beam & per fiber
- ✓ Physical Propagation (PROPER 1024² gridsize) to Lenslet Array in Pupil Plane
- ✓ Compute injection in each of 1027 fibers
- ✓ Repeat 10 times using temporal PSD per Zernike (white PSD assumed for now)



### Single wavelength as a function of time

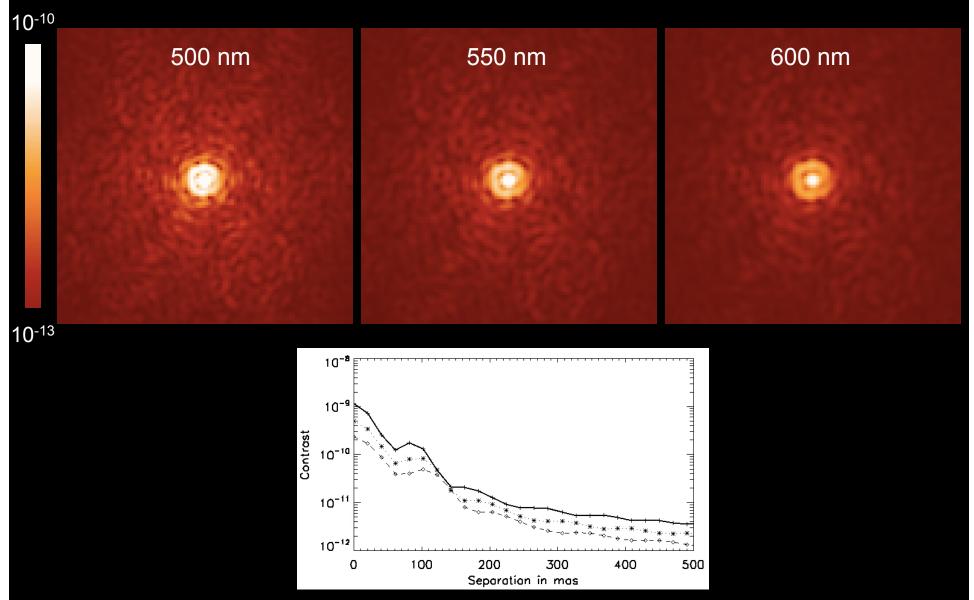
Speckle fields (central 2" x 2") at 550 nm





#### Constrast vs Wavelength

2"x2 " central speckle field (averaged over 10 instances):





### Future Improvements to Fidelity

- ✓ Input realistic spatial/ temporal distortions at telescope level etc (from thermal/ mechanical modeling)
- ✓ Include realistic input from WFC/WFS systems
- ✓ Include Nulling BC Chromatic effects (also coupled to OPD fluctuations)
- ✓ Include all optical surfaces
- ✓ Include polarization effects
- ✓ Simulate SM fiber / lenslet array imperfections (presently assumed perfect but for  $\lambda/25$  rms fiber length rms)
- ✓ Include stellar finite size



#### **Current Status**

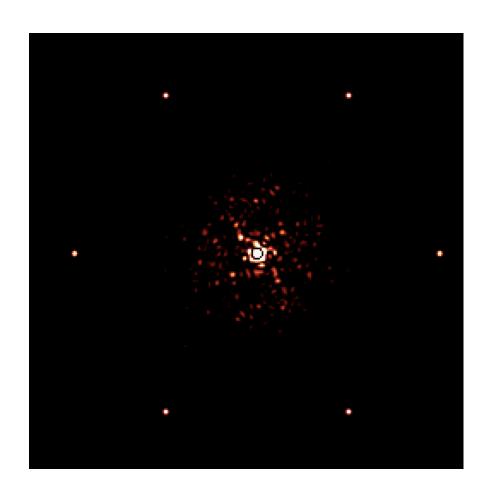
- ✓ Main Simulation "building blocks" are in:
  - Physical Optics Propagation (PROPER) from entrance aperture to final focal plane (uses angular spectrum and/or Fresnel approximation)
  - Computes SM fibers injection (both amplitude and phase effects)
  - Folds in residual spatial and temporal wavefront distortions, specifying their PSDs (ad-hoc for now, should come from more realistic inputs in terms of perturbations, WFS and WFC performance)
- ✓ Can be used to provide top-level requirements in terms of phase and amplitude stability (WFC/WFS specs)



### Back-up Slides



#### Replicas due to stepwise wavefront



With 1027 fibers and a 2m diameter sub-beams, replicas are farther than 1" in the visible